

# Non-Gaussian Data Assimilation with Stochastic PDEs: Visualizing Probability Densities of Ocean Fields?

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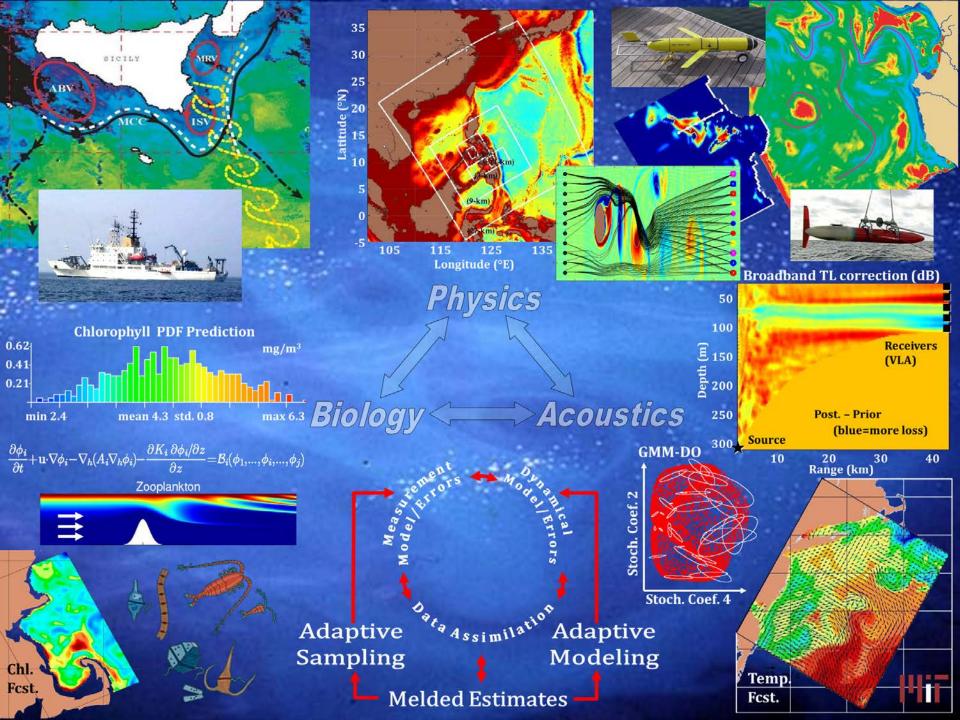
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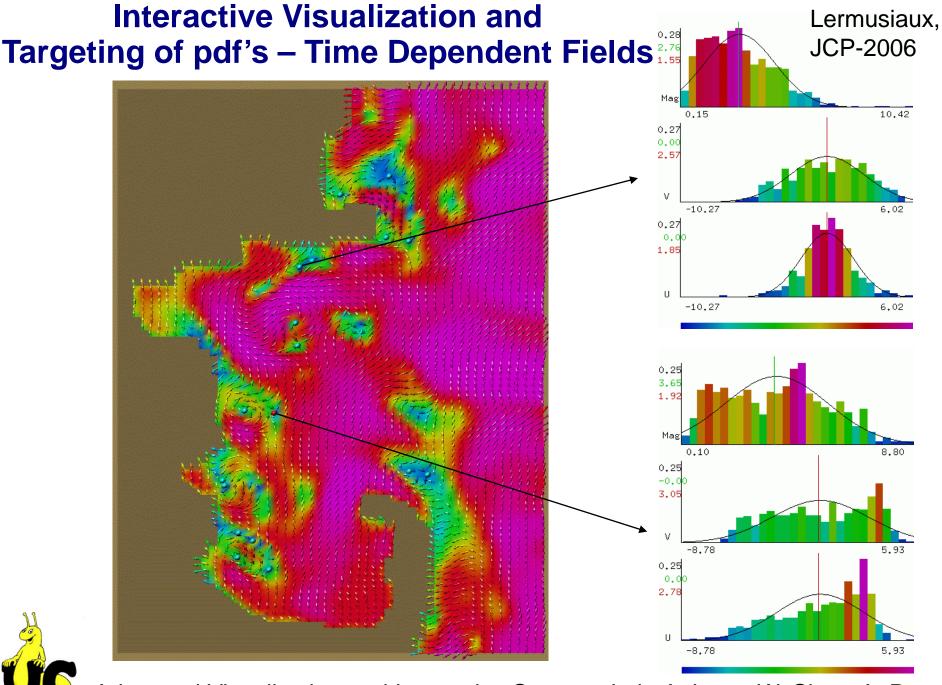
- Introduction
- Grand Challenges in Ocean/Earth-System Sciences & Engineering
  - Prognostic Equations for Stochastic Fields of Large-Dimension
  - Non-Gaussian Data Assimilation (here with DO eqns and GMM-algorithm)
- Conclusions



### Peter Muller and Frank Henyey, 1997. Workshop Assesses Monte-Carlo Simulations in Oceanography

- Ceanographers enthusiastically integrate global ocean circulation models in conjunction with atmospheric models over periods of thousands of years in order to asses future climate states – without actually knowing the skill of their ocean models …"
- \* "The case was made at the workshop that ... randomness be included in the dynamical equations .... "
- \* "... an oceanic circulation model is obtained by averaging and approximating the Navier-Stokes equations .... sub-grid-scales cannot be parameterized in terms of local mean flow quantities ... Thus, the oceanic general circulation should be regarded as a stochastic problem described by a set of stochastic PDEs."
- \* " … the vast majority of the data assimilation schemes … were derived and validated for linear systems with Gaussian noise …. The nonlinearity might actually lessen the dimensionality problem since the motion of the system might become confined … to some lower-dimensional subset of the full state space …"





Advanced Visualization and Interactive Systems Lab: A. Love, W. Shen, A. Pang

SANTA CRUZ



Mean 2.7

Mean

Mean 4.6

Mean 4.6

mg/m^3

Max 3.2

mg/m^3

Max 6.3

mg/m^3

Max 5.3

mg/m^3

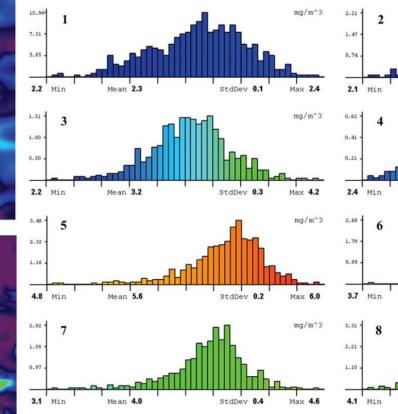
Max 5.1

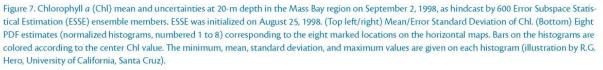
StdDev 0.2

StdDev

StdDev 0.2

StdDev 0.2





Lermusiaux et al, Oceanography-2006

Mean Chl at 7.84 20 m (mg/m3) 7.23 6.63 6.03 5.43 4.82 4.22 3.62 3.01 6 2.41 1.81 1.21 0.60 0.00 0.86 Error Std. Dev. of 0.80 Chl (mg/m3) at 20 m 0.74 0.68 0.61 0.55 0.49 0.43 0.37 0.31

8.44

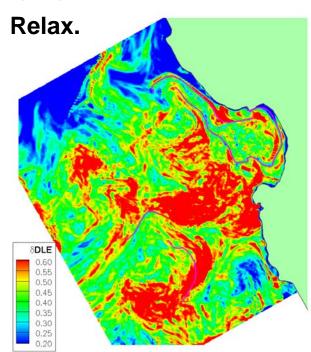
0.25

0.18

0.06

### Flow Skeletons and Uncertainties: Mean LCS overlaid on DLE error std estimate for 3 dynamical events

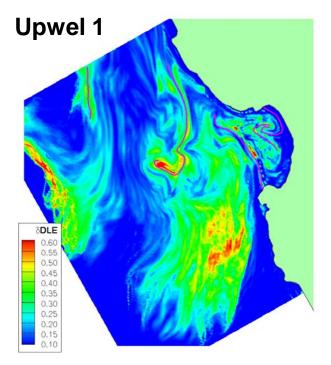
- Two upwellings and one relaxation (about 1 week apart each)
- Uncertainty estimates allow to identify most robust LCS (more intense DLE ridges are usually relatively more certain)
- Different oceanic regimes have different LCS uncertainty fields and properties

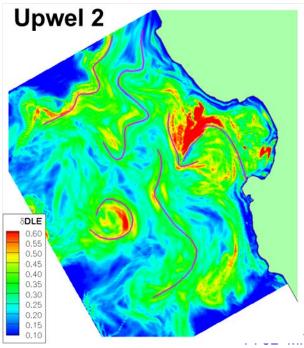


[Lermusiaux and Lekien, 2005. and In Prep, 2011

Lermusiaux, JCP-2006

Lermusiaux, Ocean.-2006]







# **Quantitatively estimate the accuracy of predictions**

### **Computational challenges for the deterministic (ocean) problem**

- Large dimensionality of the problem, <u>un-stationary</u> statistics
- Wide range of temporal and spatial scales (turbulent to climate)
- Multiple instabilities internal to the system
- Very limited observations

### Need for stochastic modeling ...

- Approximations in deterministic models including parametric uncertainties
- Initial and Boundary conditions uncertainties
- Measurement models

#### Need for data assimilation ...

- Evolve the nonlinear, i.e. non-Gaussian, correlation structures
- Nonlinear Bayesian Estimation

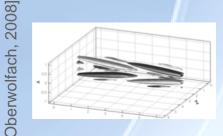


[C. Rowley,

# **Overview of Uncertainty Predictions Schemes**

$$\boldsymbol{u}(\boldsymbol{x},t;\boldsymbol{\omega}) = \boldsymbol{\overline{u}}(\boldsymbol{x},t) + \sum_{i=1}^{s} \boldsymbol{Y}_{i}(t;\boldsymbol{\omega}) \boldsymbol{u}_{i}(\boldsymbol{x},t)$$

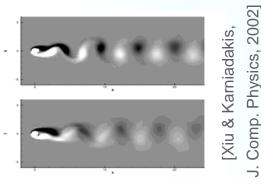
#### **Uncertainty propagation via POD method**

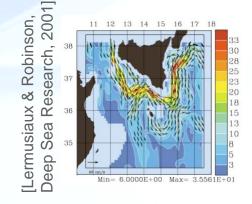


According to Lumley (Stochastic tools in Turbulence, 1971) it was introduced independently by numerous people at different times, including Kosambi (1943), Loeve (1945), Karhunen (1946), Pougachev (1953), Obukhov (1954).

#### Uncertainty propagation via generalized Polynomial-Chaos Method

Xiu & Karniadakis, J. Comp. Physics, 2002 Knio & Le Maitre, Fluid Dyn. Research, 2006 Meecham & Siegel, Phys. Fluids, 1964





Uncertainty propagation via Monte Carlo method restricted to an "evolving uncertainty subspace"

(Error Subspace Statistical Estimation - ESSE)

Lermusiaux & Robinson, MWR-1999, Deep Sea Research-2001 Lermusiaux, J. Comp. Phys., 2006

B. Ganapathysubramanian & N. Zabaras, J. Comp. Phys., 2009



# **Problem Setup: Derive equations for UQ**

#### **Statement of the problem: A Stochastic PDE**

$$\frac{\partial u(x,t;\omega)}{\partial t} = \mathcal{L}\left[u(x,t;\omega);\omega\right] \qquad x \in D$$
$$u(x,t_0;\omega) = u_0(x;\omega) \qquad x \in D \qquad \mathcal{B}\left[u/_{\partial D}\right] = h\left[\partial D;\omega\right]$$

 $\mathcal{L}[\cdot; \omega]$ Nonlinear differential operator (possibly with stochastic coefficients) $u_0(x; \omega)$ Stochastic initial conditions (given full probabilistic information) $h[\partial D; \omega]$ Stochastic boundary conditions (given full probabilistic information)

#### Goal: Evolve the full probabilistic information describing $u(x,t;\omega)$

An important representation property for the solution: Compactness  $u(x,t;\omega) = \overline{u}(x,t) + \sum_{i=1}^{s} Y_i(t;\omega) u_i(x,t)$  Advantage: Finite Dimension Evolving Subspace Disadvantage: Redundancy of representation



**Major Challenge : Redundancy**  $\boldsymbol{u}(\boldsymbol{x},t;\boldsymbol{\omega}) = \overline{\boldsymbol{u}}(\boldsymbol{x},t) + \sum_{i=1}^{n} Y_i(t;\boldsymbol{\omega}) \boldsymbol{u}_i(\boldsymbol{x},t)$ First Step (easy): Separate deterministic from stochastic/error subspace Commonly used approach: Assume that  $Y_i(t;\omega) = 0$ Second step (tricky): Evolving the finite dimensional subspace  $\mathcal{V}_s$ A separation of roles: What can  $\frac{dY_i(t;\omega)}{dt}$  tell us ? **Only** how the stochasticity evolves inside  $V_s$ source of A separation of roles: What can  $\frac{\partial u_i(x,t)}{\partial t}$  tell us ? redundancy How the stochasticity evolves **both** inside and normal to  $\mathcal{V}_s$ Natural constraint to overcome redundancy Restrict "evolution of  $\mathcal{V}_s$ " to be "normal to  $\mathcal{V}_s$ " i.e.  $\int \frac{\partial \boldsymbol{u}_i(\boldsymbol{x},t)}{\partial \boldsymbol{u}_j(\boldsymbol{x},t)} \boldsymbol{u}_j(\boldsymbol{x},t) d\boldsymbol{x} = 0 \quad \text{for all} \quad i = 1,...,s \quad \text{and} \quad j = 1,...,s$ 



# **Dynamically Orthogonal Evolution Equations**

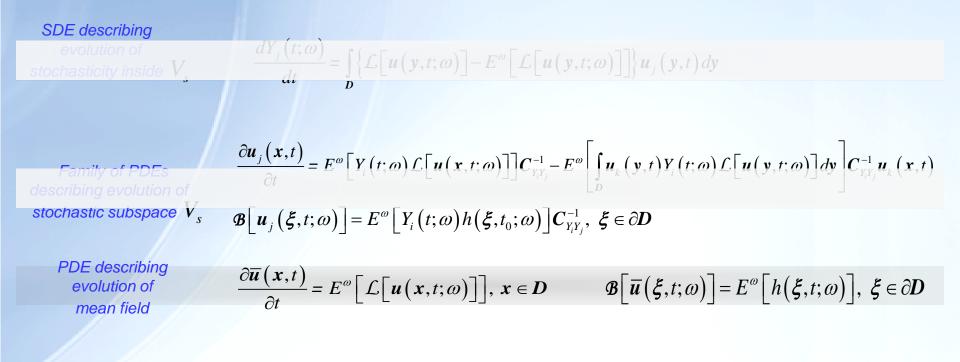
**Theorem 1:** For a stochastic field described by the evolution equation  $\frac{\partial u(x,t;\omega)}{\partial t} = \mathcal{L}\left[u(x,t;\omega);\omega\right] , \quad x \in D$   $u(x,t_0;\omega) = u_0(x;\omega) , \quad x \in D \qquad \mathcal{B}\left[u(\xi,t;\omega)\right] = h(\xi,t;\omega) , \quad \xi \in \partial D$ assuming a response of the form  $u(x,t;\omega) = \overline{u}(x,t) + \sum_{i=1}^{s} Y(t;\omega)u(x,t)$ 

assuming a response of the form  $u(x,t;\omega) = \overline{u}(x,t) + \sum_{i=1}^{3} Y_i(t;\omega) u_i(x,t)$ we obtain the following evolution equations

#### Sapsis and Lermusiaux, Physica D (2009, 2011)



## **POD & PC methods from DO equations**



Choosing a priori the stochastic subspace  $V_s$  using POD methodology we recover POD equations.

Choosing a priori the statistical characteristics of the stochastic coefficients  $Y_j(t;\omega)$  we recover the PC equations.



## **Application I : Navier-Stokes in a cavity**

2D viscous flow with stochastic initial conditions and no stochastic excitation u = U, v = 0

$$\frac{\partial u}{\partial t} + \frac{\partial P}{\partial x} = \frac{1}{\text{Re}} \Delta u - \frac{\partial (u^2)}{\partial x} - \frac{\partial (uv)}{\partial y}$$

$$\frac{\partial v}{\partial t} + \frac{\partial P}{\partial y} = \frac{1}{\text{Re}} \Delta v - \frac{\partial (uv)}{\partial x} - \frac{\partial (v^2)}{\partial y}$$

$$u = 0$$

$$v = 0$$

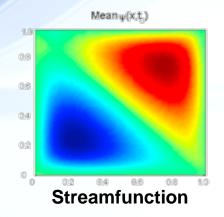
$$v = 0$$

$$v = 0$$

$$u = 0, v = 0$$

$$u = 0, v = 0$$

Initial mean flow



Initial Covariance function

$$C(r) = \left(1 + br + \frac{b^2 r^2}{3}\right) e^{-br} \quad r = ||\mathbf{x} - \mathbf{y}||$$
  

$$\int C(||\mathbf{x} - \mathbf{y}||) \hat{u}_i(\mathbf{x}) d\mathbf{x} = \lambda_i^2 \hat{u}_i(\mathbf{y})$$
  

$$u_{0,i}(\mathbf{x}) = \hat{u}_i(\mathbf{x})$$
  

$$Y_i(t_0; \omega) \sim \mathcal{N}(0, \lambda_i)$$

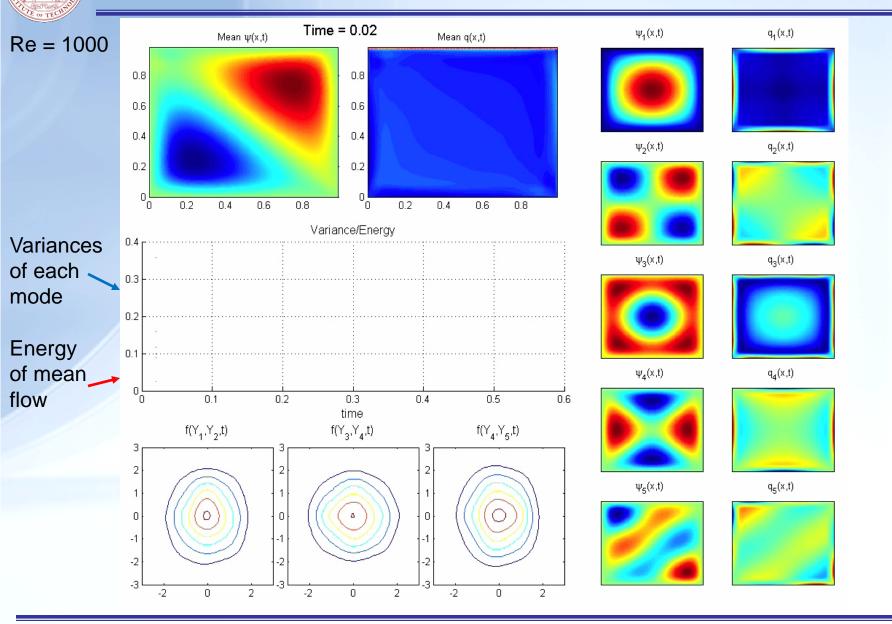
PDE Numerics: C-grid, upwind [M. Griebel et al., 1998]

SDE Numerics: here, s-dimensional Monte-Carlo

 $\psi_{i}(\mathbf{x},\mathbf{\xi})$ 

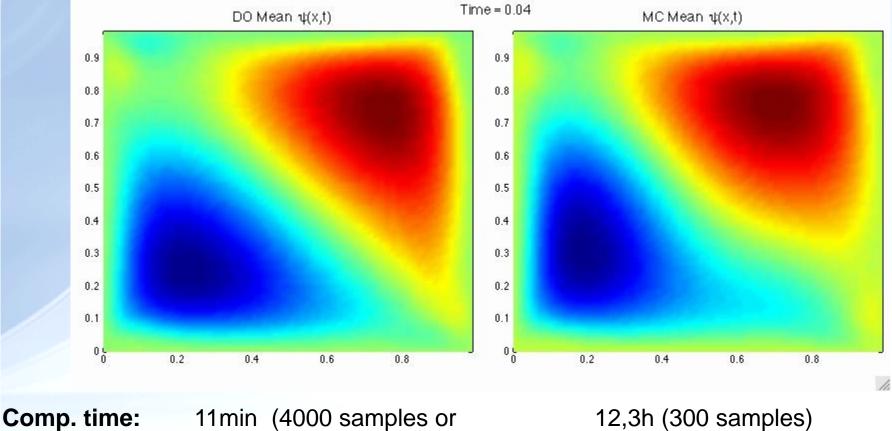
 $\psi_{2}(\mathbf{x},\mathbf{t})$ 

## **Application I : Navier-Stokes in a cavity**



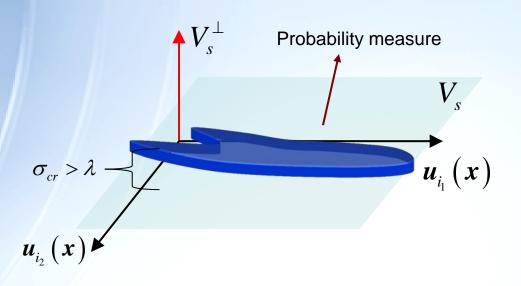


### **Comparison with Monte-Carlo**



11min (4000 samples or analytical  $Y_i$ )

### Adapt the stochastic subspace dimension



- In the context of DO equations so far the size of the stochastic subspace V<sub>s</sub> remained invariant.
- For intermittent or transient phenomena the dimension of the stochastic subspace may vary significantly with time. This is accounted for by ESSE.

We need criteria to evolve the dimensionality of the stochastic subspace

This is a particularly important issue for stochastic systems with deterministic initial conditions



#### **Dimension Reduction**

Comparison of the minimum eigenvalue of the correlation matrix  $C_{Y_iY_i}$ .

$$\lambda_{\min}\left[C_{Y_iY_j}\right] < \sigma_{cr} \longrightarrow$$
 pre-defined value

Removal of the corresponding direction from the stochastic subspace.

#### **Dimension Increase**

Comparison of the minimum eigenvalue of the correlation matrix  $C_{Y_iY_i}$ .

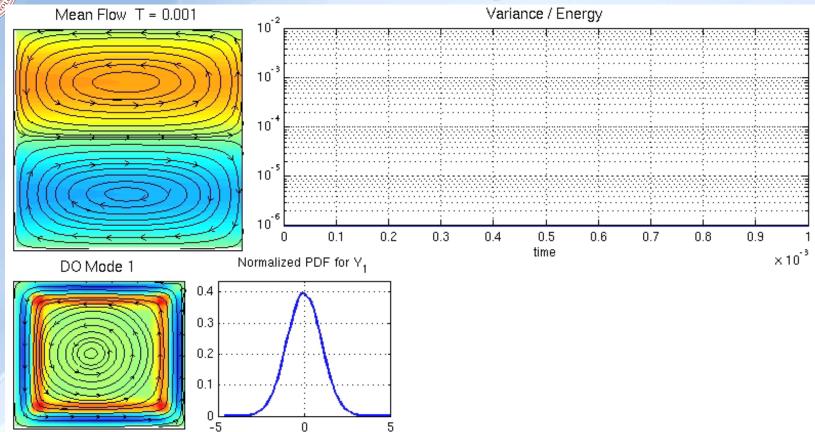
$$\lambda_{\min}\left[\boldsymbol{C}_{Y_iY_j}\right] > \Sigma_{cr} \longrightarrow \text{pre-defined value}$$

Addition of a new direction  $\boldsymbol{u}_i(\boldsymbol{x},t)$  in the stochastic subspace  $V_s$ .

#### How do we choose this new direction? By breeding in the orthogonal complement $V_s^{\perp}$

Same problem when we start with deterministic initial condition (dimension of stochastic subspace is zero)

## Example: Double Gyre, Re=10,000

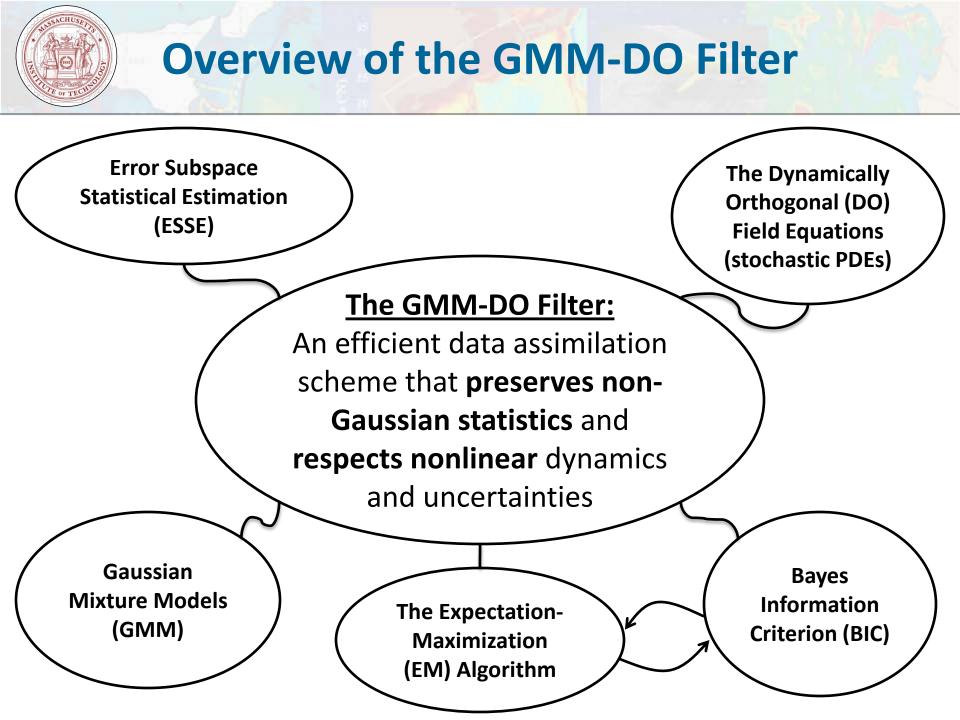




# **The GMM-DO Filter:**

# Data Assimilation and Adaptive Sampling with Gaussian Mixture Models using the Dynamically Orthogonal field equations

(Sondergaard, 2011; Sondergaard and Lermusiaux, MWR-to-be-submitted, Parts I and II)





# Gaussian Mixture Models (with Bayesian update)

The probability density function for a random vector, **x**, distributed according to a multivariate Gaussian mixture model is given by

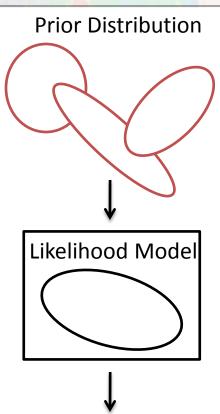
$$p_{\boldsymbol{X}}(\boldsymbol{x}) = \sum_{j=1}^{M} \pi_{j} \times N(\boldsymbol{x}; \boldsymbol{\mu}_{j}, \boldsymbol{\Sigma}_{j}),$$

subject to the constraint that

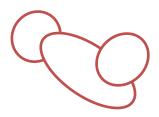
$$\sum_{j=1}^{M} \pi_j = 1.$$

We refer to M as the mixture complexity and  $\pi_j$  as the mixture weights. The multivariate Gaussian density function takes the form:

$$N(x;\mu,\Sigma) = \frac{1}{(2\pi)^{n/2} |\Sigma|^{1/2}} e^{-\frac{1}{2}(x-\mu)^T \Sigma^{-1}(x-\mu)}.$$



**Posterior Distribution** 







Kernel Density Approximation



# **Overview of the GMM-DO Filter**

**Initial Conditions:** Initialize the state vector in a decomposed form that accords with the Dynamically Orthogonal field equations:

$$x_{r,0} = \bar{x}_0 + \mathcal{X}_0 \phi_{r,0}, \quad r = \{1, \dots, N\}.$$

 $\bar{\boldsymbol{x}} \in \mathbb{R}^n$  is the mean vector,  $\boldsymbol{\mathcal{X}} \in \mathbb{R}^{n \times s}$  defines the matrix of modes describing an orthonormal basis for the stochastic subspace, and the  $\phi_r \in \mathbb{R}^s$  represent N realizations drawn from the multivariate random vector described by  $\{\Phi_{1,0}(\omega), \ldots, \Phi_{s,0}(\omega)\}$  that reside in the stochastic subspace of dimension s.

**Forecast:** Using either the initial DO conditions or the posterior state description following the assimilation of data at time k - 1,

 $\boldsymbol{x}_{r,k-1}^{a} = \bar{\boldsymbol{x}}_{k-1}^{a} + \boldsymbol{\mathcal{X}}_{k-1} \boldsymbol{\phi}_{r,k-1}^{a}, \quad r = \{1, \dots, N\},$ 

apply the DO equations to efficiently evolve the probabilistic description of the state vector in time, arriving at a forecast for observation time k:

$$\boldsymbol{x}_{r,k}^f = \bar{\boldsymbol{x}}_k^f + \boldsymbol{\mathcal{X}}_k \boldsymbol{\phi}_{r,k}^f, \quad r = \{1, \dots, N\}.$$

**Observation:** Employ a linear (or linearized) observation model,

$$\boldsymbol{Y}_k = \boldsymbol{H} \boldsymbol{X}_k + \boldsymbol{\Upsilon}_k, \quad \boldsymbol{\Upsilon}_k \sim \mathcal{N}\left(\boldsymbol{\upsilon}_k; \boldsymbol{0}, \boldsymbol{R}\right).$$

where  $\mathbf{Y}_k \in \mathbb{R}^p$  is the observation random vector at discrete time k;  $\mathbf{H} \in \mathbb{R}^{p \times n}$  is the linear observation model; and  $\mathbf{\Upsilon}_k \in \mathbb{R}^p$  the corresponding random noise vector, assumed to be of a Gaussian distribution. The realized observation vector is denoted by  $\mathbf{y}_k \in \mathbb{R}^p$ .

(Sondergaard, 2011; Sondergaard and Lermusiaux, MWR-to-be-submitted, Parts I and II)

### **GMM Filter Example:**

### Flow Exiting a Strait or "Sudden Expansion Flow"

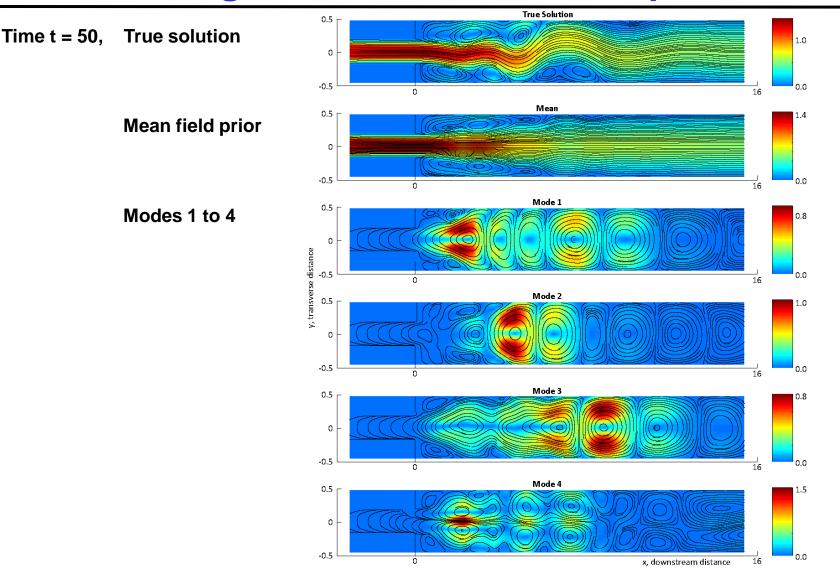


Figure 5-15: True solution; DO mean field; and first four DO modes at the first assimilation step, time T = 50.



# **Overview of the GMM-DO Filter**

#### **GMM fit in DO stochastic subspace**

**Forecast:** Using either the initial DO conditions or the posterior state description following the assimilation of data at time k - 1,

$$\boldsymbol{x}_{r,k-1}^{a} = \bar{\boldsymbol{x}}_{k-1}^{a} + \boldsymbol{\mathcal{X}}_{k-1} \boldsymbol{\phi}_{r,k-1}^{a}, \quad r = \{1, \dots, N\},$$

apply the DO equations to efficiently evolve the probabilistic description of the state vector in time, arriving at a forecast for observation time k:

$$\boldsymbol{x}_{r,k}^f = \bar{\boldsymbol{x}}_k^f + \boldsymbol{\mathcal{X}}_k \boldsymbol{\phi}_{r,k}^f, \quad r = \{1, \dots, N\}.$$

**Fitting of GMM:** For Gaussian mixture models of increasing complexity (i.e. M = 1, 2, 3, ...), repeat until a minimum of the BIC is met:

*i.* Use the EM algorithm to obtain the prior mixture parameters

$$\pi_{j,k}^f, \boldsymbol{\mu}_{j,k}^f, \boldsymbol{\Sigma}_{j,k}^f, \quad j = 1, \dots, M$$

within the stochastic subspace based on the set of ensemble realizations,  $\{\phi_k^f\} = \{\phi_{1,k}^f, \dots, \phi_{N,k}^f\}.$ 

*ii.* Use the Bayesian Information Criterion to evaluate the fit of the Gaussian mixture model for the given M.

(Sondergaard, 2011; Sondergaard and Lermusiaux, MWR-to-be-submitted, Parts I and II)



# The EM algorithm with GMM

Based on the data at hand, the Expectation-Maximization algorithm describes an iterative procedure for obtaining the **Maximum Likelihood** estimate for the unknown set of parameters,  $\theta$ , here of our prior Gaussian mixture model:

$$\{\pi_1,\ldots,\pi_M,\mu_1,\ldots,\mu_M,\Sigma_1,\ldots,\Sigma_M\}.$$

**Procedure.** Given the n data, x, and initial parameter estimate  $\theta(0)$ , repeat until convergence:

(1) Expectation: Using the current set of parameters,  $\theta(k)$ , form

$$\tau_j(\boldsymbol{x}_i; \boldsymbol{\theta}^{(k)}) = \frac{\pi_j^{(k)} N\left(\boldsymbol{x}_i; \boldsymbol{\mu}_j^{(k)}, \boldsymbol{\Sigma}_j^{(k)}\right)}{\sum_{m=1}^M \pi_k^{(k)} N\left(\boldsymbol{x}_i; \boldsymbol{\mu}_m^{(k)}, \boldsymbol{\Sigma}_m^{(k)}\right)}$$

(2) Minimization: Update the estimate for the set of parameters,  $\theta(k+1)$ , according to

$$\begin{aligned} \pi_j^{(k+1)} &= \frac{\sum_{i=1}^n \tau_j \left( x_i; \theta^{(k)} \right)}{n} = \frac{n_j^{(k)}}{n} \\ \mu_j^{(k+1)} &= \frac{1}{n_j^{(k)}} \sum_{i=1}^n \tau_j \left( x_i; \theta^{(k)} \right) x_i \\ \Sigma_j^{(k+1)} &= \frac{1}{n_j^{(k)}} \sum_{i=1}^n \tau_j \left( x_i; \theta^{(k)} \right) \left( x_i - \mu_j^{(k)} \right) \left( x_i - \mu_j^{(k)} \right)^T. \end{aligned}$$

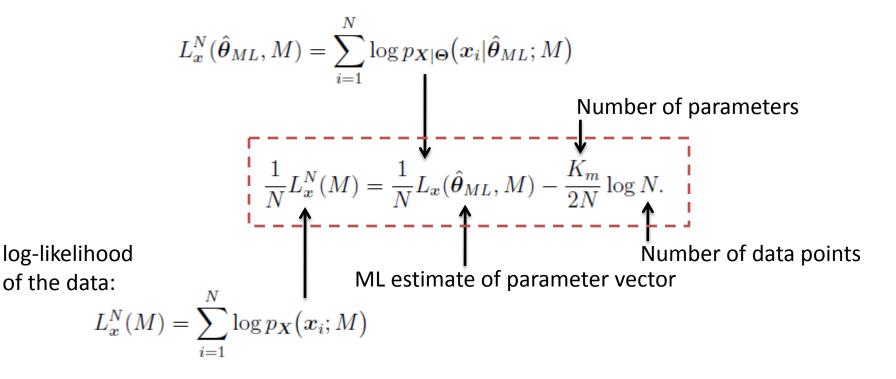


# **Bayes Information Criterion**

Determining the complexity of a Gaussian mixture model can be put in the context of **model selection**: based on the data at hand, **x**, we wish to select the model complexity that maximizes the likelihood of this data:

$$p_{X}(x; M) = \int p_{X|\Theta}(x|\theta; M) p_{\Theta}(\theta; M) d\theta$$

We use **Bayes Information Criterion** -- we select the simplest hypothesis consistent with the data, i.e. maximize the log-likelihood of the data around the EM-ML estimate of the parameters:



### GMM Filter Example: Flow Exiting a Strait or "Sudden Expansion Flow"

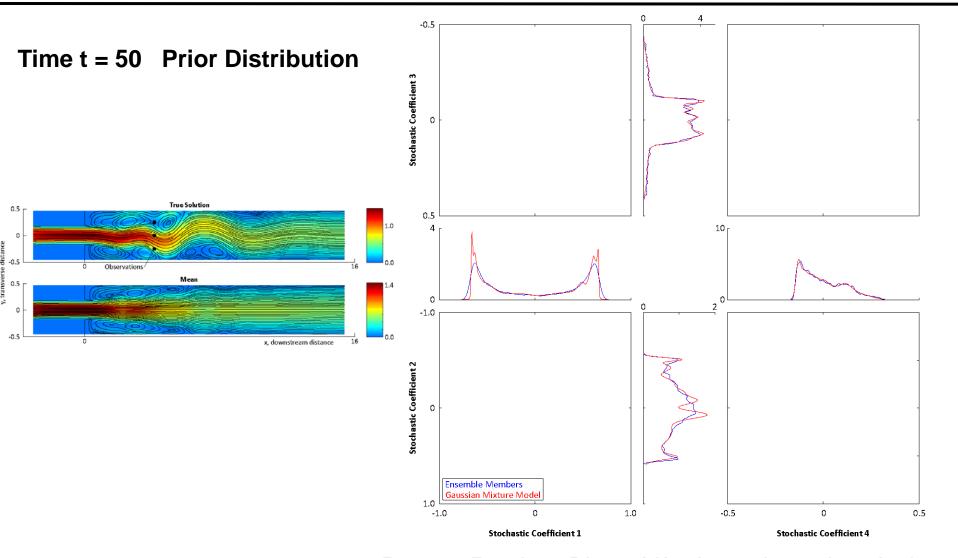


Figure 5-16: True solution; DO mean field; and joint and marginal prior distribution, identified by the Gaussian mixture model of complexity 29, and associated ensembles of the first four modes at the first assimilation step, time T = 50.

## GMM Filter Example: Flow Exiting a Strait or "Sudden Expansion Flow"

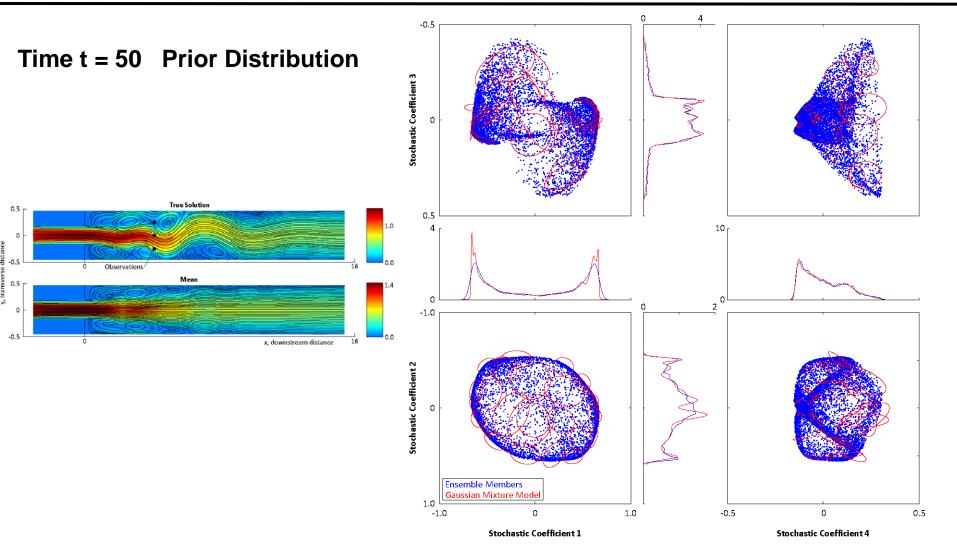


Figure 5-16: True solution; DO mean field; and joint and marginal prior distribution, identified by the Gaussian mixture model of complexity 29, and associated ensembles of the first four modes at the first assimilation step, time T = 50.

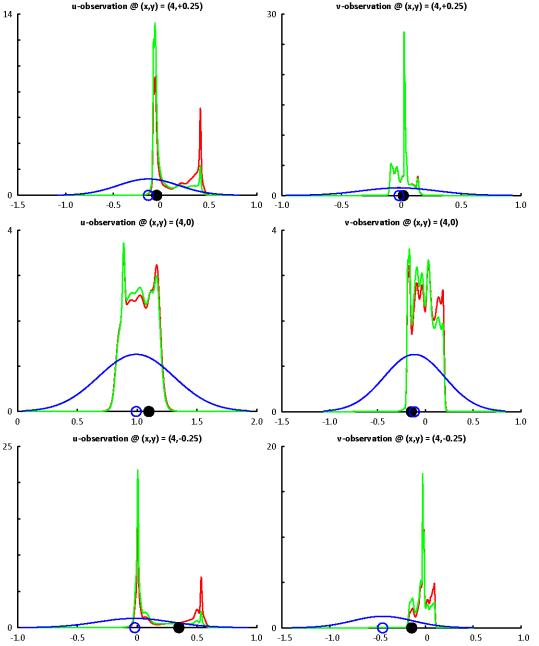
#### GMM Filter Example: Flow Exiting a Strait or "Sudden Expansion Flow" <sup>14</sup>

Time t = 50

**Observations and their pdf** 

**Prior Distributions at these data points** 

Posterior Distributions at these data points



Legend: Observation Observation Obstribution Prior Distribution Posterior Distribution

Figure 5-17: True solution; observation and its associated Gaussian distribution; and the prior and posterior distributions at the observation locations at time T = 50.



### **Overview of the GMM-DO Filter**

**Observation:** Employ a linear (or linearized) observation model,

 $oldsymbol{Y}_k = oldsymbol{H}oldsymbol{X}_k + oldsymbol{\Upsilon}_k, \quad oldsymbol{\Upsilon}_k \sim \mathcal{N}\left(oldsymbol{v}_k; oldsymbol{0}, oldsymbol{R}
ight).$ 

where  $\mathbf{Y}_k \in \mathbb{R}^p$  is the observation random vector at discrete time k;  $\mathbf{H} \in \mathbb{R}^{p \times n}$  is the linear observation model; and  $\mathbf{\Upsilon}_k \in \mathbb{R}^p$  the corresponding random noise vector, assumed to be of a Gaussian distribution. The realized observation vector is denoted by  $\mathbf{y}_k \in \mathbb{R}^p$ .

### Bayesian Update of GMM in DO stochastic subspace

For GMM-DO Update Theorem, see:

(Sondergaard and Lermusiaux, MWR -to-be-submitted -2011, Parts I and II) Update:

*i.* Compute:

$$egin{aligned} ilde{m{H}}_k &\equiv m{H}m{\mathcal{X}}_k \ ilde{m{y}}_k &\equiv m{y}_k - m{H}ar{m{x}}_k^f \end{aligned}$$

and determine the individual Kalman gain matrices:

$$\tilde{\boldsymbol{K}}_{j,k} = \boldsymbol{\Sigma}_{j,k}^{f} \tilde{\boldsymbol{H}}_{k}^{T} (\tilde{\boldsymbol{H}}_{k} \boldsymbol{\Sigma}_{j,k}^{f} \tilde{\boldsymbol{H}}_{k}^{T} + \boldsymbol{R})^{-1}, \quad j = 1, \dots, M.$$

ii. Assimilate the measurement,  $y_k$ , by calculating the 'intermediate' mixture means in the stochastic subspace,

$$\hat{\boldsymbol{\mu}}_{j,k}^{a} = \boldsymbol{\mu}_{j,k}^{f} + \tilde{\boldsymbol{K}}_{j,k}(\tilde{\boldsymbol{y}}_{k} - \tilde{\boldsymbol{H}}_{k}\boldsymbol{\mu}_{j,k}^{f}),$$

and further compute the posterior mixture weights:

$$\pi_{j,k}^{a} = \frac{\pi_{j,k}^{f} \times \mathcal{N}\left(\tilde{\boldsymbol{y}}_{k}; \tilde{\boldsymbol{H}}_{k} \boldsymbol{\mu}_{j,k}^{f}, \tilde{\boldsymbol{H}}_{k} \boldsymbol{\Sigma}_{j,k}^{f} \tilde{\boldsymbol{H}}_{k}^{T} + \boldsymbol{R}\right)}{\sum_{l=1}^{M} \pi_{l,k}^{f} \times \mathcal{N}\left(\tilde{\boldsymbol{y}}_{k}; \tilde{\boldsymbol{H}}_{k} \boldsymbol{\mu}_{l,k}^{f}, \tilde{\boldsymbol{H}}_{k} \boldsymbol{\Sigma}_{l,k}^{f} \tilde{\boldsymbol{H}}_{k}^{T} + \boldsymbol{R}\right)}$$

iii. Update the DO mean field,

$$ar{oldsymbol{x}}_k^a = ar{oldsymbol{x}}_k^f + oldsymbol{\mathcal{X}}_k \sum_{j=1}^M \pi_{j,k}^a imes oldsymbol{\hat{\mu}}_{j,k}^a,$$

as well as the mixture parameters within the stochastic subspace:

$$\boldsymbol{\mu}_{j,k}^{a} = \hat{\boldsymbol{\mu}}_{j,k}^{a} - \sum_{j=1}^{M} \pi_{j,k}^{a} \times \hat{\boldsymbol{\mu}}_{j,k}^{a}$$
$$\boldsymbol{\Sigma}_{j,k}^{a} = (\boldsymbol{I} - \tilde{\boldsymbol{K}}_{j,k} \tilde{\boldsymbol{H}}_{k}) \boldsymbol{\Sigma}_{j,k}^{f}.$$

iv. Generate the posterior set of subspace ensemble realizations,  $\{\phi_k^a\} = \{\phi_{1,k}^a, \dots, \phi_{N,k}^a\}$ , according to the multivariate Gaussian mixture model with posterior parameter values:

$$\pi^a_{j,k}, \boldsymbol{\mu}^a_{j,k}, \boldsymbol{\Sigma}^a_{j,k}, \quad j = 1, \dots, M.$$

### GMM Filter Example: Flow Exiting a Strait or "Sudden Expansion Flow"

Time t = 50

#### **Posterior Distribution**

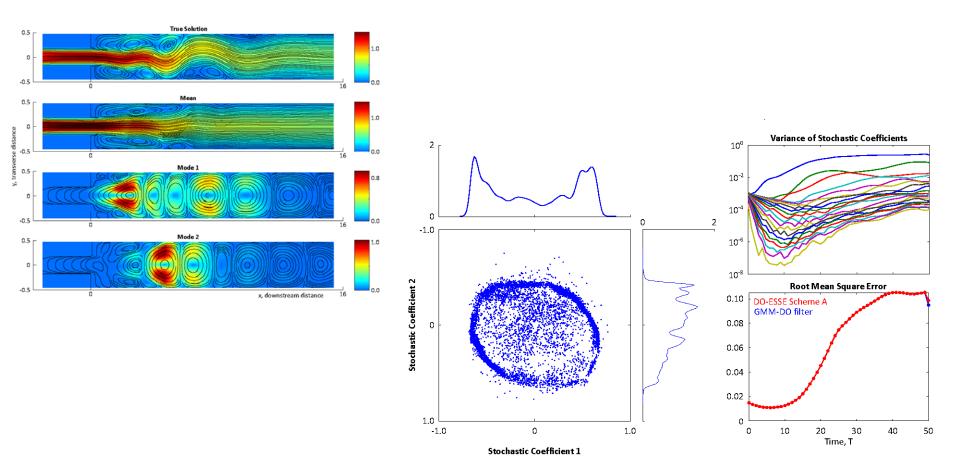


Figure 5-18: True solution; condensed representation of the posterior DO decomposition; and root mean square errors at time T = 50.

### **GMM-DO Filter:**

## **DO equations and Non-Gaussian Data Assimilation**

"Flow exiting a Strait" Test Case: Results show that our new DO equations and Non-Gaussian assimilation leads to optimal error reduction

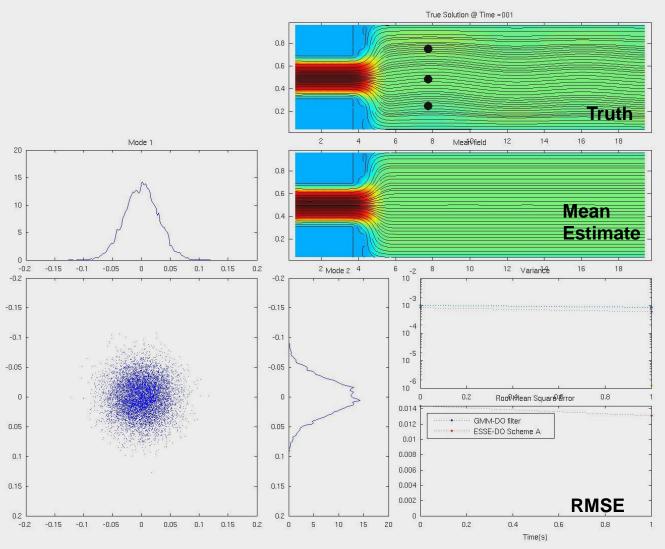
(Top Right): True solution mean flow-field streamlines overlaid on vorticity with sampling positions as circles

(Left): DO marginal pdfs represented as samples with the single 1<sup>st</sup> and 2<sup>nd</sup> DO marginal pdfs on each sides, clearly showing non-Gaussian behavior.

(Middle Right): Mean estimate mean using GMM-DO filter.

(Bottom Right): Variance of 10 DO modes as a function of time

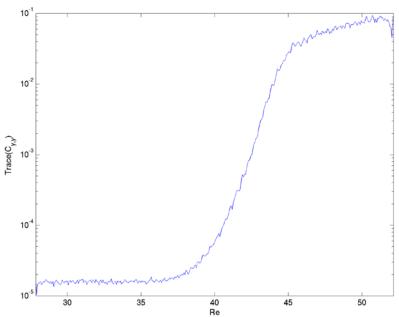
(Bottom Right): Comparisons of the root-mean-square-error (truth minus mean) as a function of time, clearly showing superior performance of GMM-DO filter

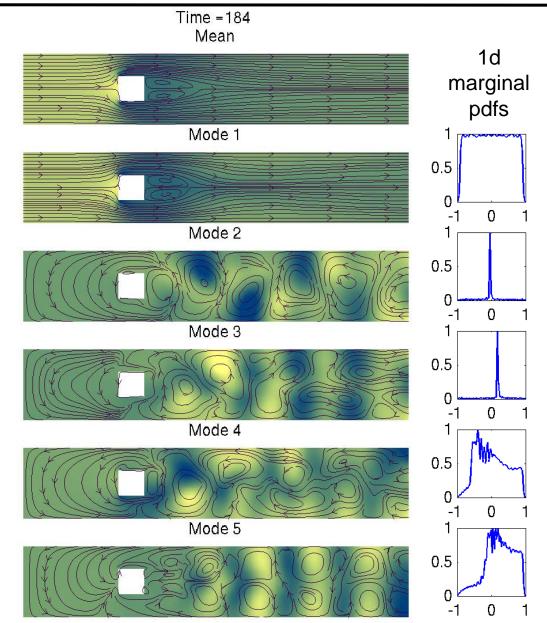


## **Visualizing Uncertainty in Fluid and Ocean Flows**

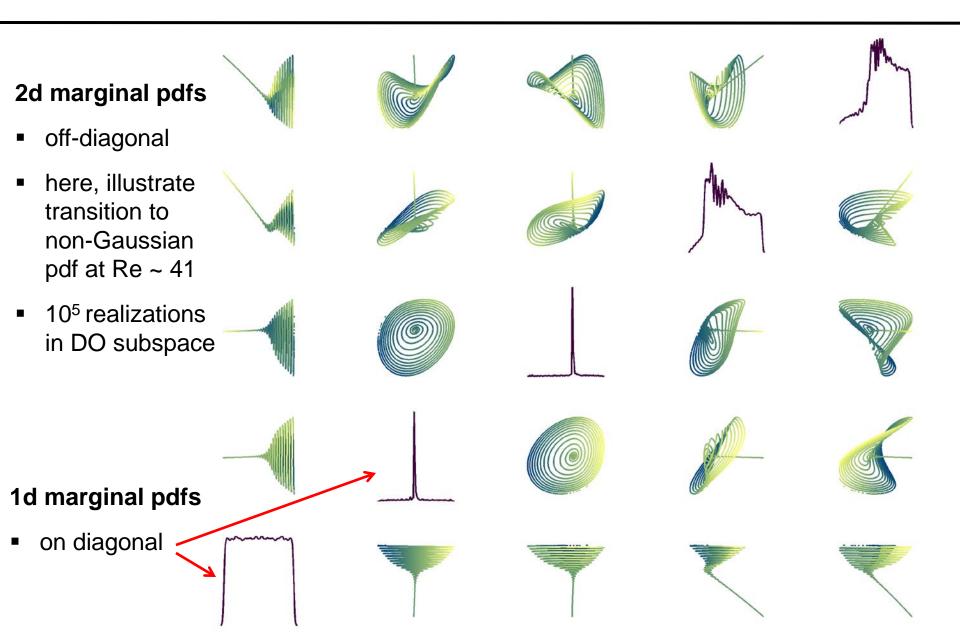
# Stochastic Flow behind a square cylinder

- Uncertain initial and boundary conditions
- Range of Reynolds number modeled with a single DO simulation
- Equivalent to 10<sup>5</sup> deterministic runs





## **Visualizing Uncertainty in Fluid and Ocean Flows**



### **Visualizing Uncertainty in Fluid and Ocean Flows**

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2d marginal pdfs	-	Á		4	٩		ŀ	L	-
<ul> <li>9 DO modes</li> <li>Still 10<sup>5</sup> realizations in DO subspace</li> </ul>	<b>\</b>	4		<u>«</u>	4	1	M	4	2
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# CONCLUSIONS

### Prognostic DO Equations for Stochastic Fields

GMM-DO Data Assimilation

### Visualizing Probability Densities of Ocean Fields?

- Scientific Visualization of Uncertainty
  - Overlays (pseud-color, contours, etc)
  - Histograms at each point in physical space, time-dependent
  - Key question: how to visualize pdfs in DO subspace?, but then in physical space?
- Societal Visualization of Uncertainty
  - Overlays
  - Direct Volume rendering, Transparency
  - Glyphs, etc